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PHOTOMETRY OF HIGH-AMPLITUDE DELTA SCUTI STARS IN 2013

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This report is the sixth paper in a series on photometry of High-Amplitude Delta Scuti (HADS) Stars (see Wils et al., 2013 for references to the earlier papers in this series). It presents details on 390 times of maximum for 75 HADS, obtained during 2013. The method used to calculate the times of maximum is described in the first paper of the series (Wils et al., 2009).

Table 1 lists the code used for the observers and their instruments. We list the observed maxima in Table 2 with star name (Col. 1), the epoch of the observed maximum (Col. 2), the uncertainty of the epoch (Col. 3), the observer's code (Col. 4) and the filter used (Col. 5) for the observations. When a maximum was observed in more than one filter by the same observer, the table shows the average value of the times obtained in each filter individually (note that there may be a significant delay between the maximum times when observed in different filters).

We found that the observed times of maximum deviated considerably from their ephemeris for some stars. Insufficient precision in the original determination of the period is often the cause. We give new elements for these stars in Table 3. To obtain the highest possible precision, we used data from the ASAS (Pojmański, 2002), NSVS (Woźniak et al., 2004), CRTS (Drake et al., 2009) and SuperWASP surveys (Butters et al., 2010) in conjunction with our own data.

We detected a sudden change in the period of V376 Cam, likely at the beginning of 2013. A plot of the $O - C$ data calculated from our observations using a linear ephemeris are given in Fig. 1. Assuming a constant period before and after the change, the period shortened by 0.282 ± 0.006 seconds, or $\Delta P/P = 2.3 \times 10^{-5}$. Sweigart and Renzini (1979) proposed small random mixing events in a semiconvective zone and at the edge of the

Table 1: List of instruments used for the observations.

Code	Observer(s)	Telescope	Observatory	CCD
AA	AA	Catadioptric 30 cm	Perseus Observatory	SBIG ST-10XME
AB	AB	Catadioptric 35 cm	Carpe Noctem Observatory	SBIG ST-9E
FN	FN	Catadioptric 40 cm	Alkmaar, Nederland	SBIG ST-7XME
FNA	FN	Catadioptric 25 cm	ABT Metius	SBIG ST402XME
HMBW	FJH	Catadioptric 30 cm	Astrokolhoz, New Mexico	SBIG ST-9XE
HMBC	FJH	Catadioptric 40 cm	Chile Remote Obs. Atacama Desert	FLI ML16803
HO40	PL+PVC	Newton 40 cm	R.O.B.-Humain	SBIG ST-10XME
IS	IS	Catadioptric 25 cm	ABT Metius	SBIG ST402XME
MAV	MV	Maksutov 26 cm	Leest Observatory	SBIG ST-10XME
MAVN	MV	Newton 35 cm	Leest Observatory	QSI583 WSG
RP	RDP	Catadioptric 36 cm	Shobdon, UK	Starlight XPress SXV-H9
SBL	BS	Cassegrain 28 + 23.5 cm	Alan Guth Observatory	Starlight XPress MX-716
SK	SK	Catadioptric 30 cm	Zagori Observatory	SBIG ST-7XMEI
SO30	CWR	Catadioptric 30 cm	SETEC Observatory	SBIG ST-8iXME
SO40	CWR	Catadioptric 40 cm	SETEC Observatory	SBIG ST-8XME
VWSR	JVW	Refractor 15.2 cm	Hooglede, Belgium	SBIG ST-7XME
VWS	JVW	Catadioptric 23.5 cm	Hooglede, Belgium	SBIG ST-8XME

convective core as the cause of sudden period changes in RR Lyrae stars. A similar mechanism may be at work here. Alternatively the light time effect from the passage through periastron of a companion in a highly elliptic orbit may mimic a sudden change in period. In that case these events should reoccur at regular intervals.

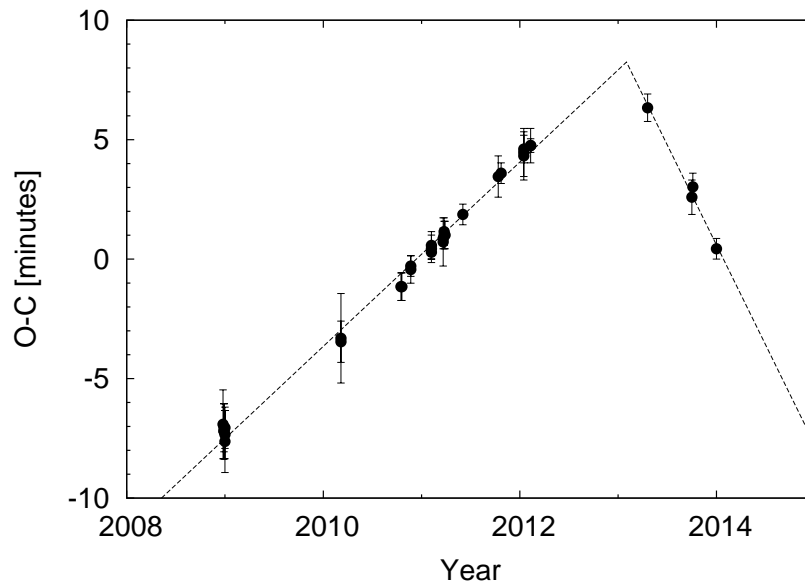


Figure 1. $O - C$ values for the maxima of V376 Cam with respect to the ephemeris $\text{Max HJD} = 2454823.4194 + 0^{\text{d}}14032268 \times E$. The straight lines indicate the instantaneous period before and after the change.

Four of the observed stars observed turned out be multiperiodic variables. We found V879 Her (observed by AA), GSC 1489-0914 (observed by AA) and GSC 4145-0919 (observed by JVW and AA) to be double-mode pulsators while GSC 1566-2802 (observed by AA and MV) showed at least one other pulsation mode, probably a non-radial mode. Table 4 lists details about the independent frequencies f_0 , the fundamental radial mode, and f_1 , the first overtone mode in the case of the double-mode pulsators and a non-radial

mode for GSC 1566-2802. The frequencies, amplitudes and phases, and their uncertainties were determined using Period04 (Lenz & Breger, 2005). The uncertainties were derived using Monte Carlo simulations. We also detected a number of linear combinations of the independent frequencies in all stars. In 2014 Hümmerich independently discovered the double-mode behaviour of GSC 1489-0914¹. Data of these four stars are available from the IBVS website: 6122-t1.txt (V879 Her); 6122-t2.txt (GSC 1489-0914); 6122-t3.txt, 6122-t4.txt (GSC 4145-0919); 6122-t5.txt, 6122-t6.txt (GSC 1566-2802).

One of us (JVW) found one of the comparison stars used for LW Dra, GSC 4431-0698, to be a W UMa type variable with a period of 0.36949 days. GSC 4431-0698 is a close pair of two stars (2MASS J19053535+6840195 and 2MASS J19053648+6840169) 6".7 apart. We could not determine which one of these stars is the variable. The data are available online (6122-t7.txt).

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¹AAVSO VSX: "Fundamental period and epoch are tabulated. First overtone period: 0.0419105 d. Period ratio: 0.7734. J-K=0.15 (2MASS); B-V=0.21 (APASS)" (<http://www.aavso.org/vsx>)

Table 2: Observed times of maximum (Epoch = HJD - 2400000).

Star	Epoch	Unc.	Obs.	Filter	Star	Epoch	Unc.	Obs.	Filter	
GP And	56573.6020	0.0005	SO30	V	V792 Cep	56460.5081	0.0012	AA	C	
	56573.6808	0.0006	SO30	V		56497.3300	0.0013	AA	C	
	56573.7592	0.0005	SO30	V		56497.4624	0.0012	AA	C	
	56573.8381	0.0005	SO30	V		XX Cyg	56487.5234	0.0003	MAV	V
	56573.9166	0.0007	SO30	V		V2455 Cyg	56482.4685	0.0004	MAV	V
	56577.6149	0.0007	SO30	V		LW Dra	56403.4120	0.0010	VWSR	V
	56577.6931	0.0005	SO30	V		56410.3829	0.0012	VWSR	V	
	56577.7718	0.0005	SO30	V		56423.4980	0.0006	VWSR	V	
	56577.8508	0.0008	SO30	V		56451.5004	0.0007	VWSR	V	
	56577.9295	0.0004	SO30	V		V1116 Her	56352.9694	0.0017	HMBW	V
	56604.6032	0.0007	SO30	V		56404.3796	0.0018	SBL	V	
	56604.6819	0.0006	SO30	V		56404.4768	0.0016	SBL	V	
	56604.7606	0.0007	SO30	V		56404.5702	0.0016	SBL	V	
	56604.8389	0.0006	SO30	V		56466.3973	0.0010	AA	C	
	56615.4615	0.0006	RP	V		56466.4922	0.0010	AA	C	
56641.6626	0.0004	SO30	V	V1209 Her	56351.9971	0.0004	HMBW	V		
V460 And	56506.5001	0.0005	MAV	V	56404.3703	0.0004	MAVN	V		
	56566.3358	0.0004	MAV	V	56404.4218	0.0004	MAVN	V		
	56566.4090	0.0009	MAV	V	56404.4728	0.0003	MAVN	V		
V524 And	56573.5811	0.0005	SO40	V	56404.5241	0.0003	MAVN	V		
	56573.6753	0.0003	SO40	V	KZ Lac	56483.4106	0.0018	MAV	V	
	56573.7698	0.0003	SO40	V	56483.5146	0.0014	MAV	V		
	56573.8644	0.0002	SO40	V	56507.4263	0.0010	MAV	V		
V528 And	56573.9590	0.0003	SO40	V	56508.4705	0.0017	MAV	V		
	56543.5202	0.0012	AA	C	56508.5747	0.0020	MAV	V		
	56590.2982	0.0013	AA	C	56541.3611	0.0004	AB	C		
	56590.3870	0.0011	AA	C	56577.5934	0.0012	SO40	V		
	56590.4761	0.0011	AA	C	56577.6979	0.0012	SO40	V		
V544 And	56581.2459	0.0012	AA	C	56577.8026	0.0015	SO40	V		
	56581.3541	0.0008	AA	C	56577.9072	0.0013	SO40	V		
	56587.2341	0.0005	AA	C	SZ Lyn	56299.6526	0.0005	SO30	V	
	56587.3408	0.0005	AA	C	56299.7733	0.0009	SO30	V		
	56587.4482	0.0004	AA	C	56299.8940	0.0007	SO30	V		
CY Aqr	56587.5549	0.0003	AA	C	56309.6568	0.0009	SO30	BV		
	56559.4172	0.0008	AB	C	56309.7770	0.0005	SO30	BV		
YZ Boo	56339.8669	0.0006	HMBW	V	56309.8979	0.0010	SO30	BV		
	56339.9708	0.0007	HMBW	V	56310.0187	0.0016	SO30	BV		
	56385.3547	0.0006	MAVN	V	56310.6211	0.0010	SO30	BV		
	56385.4589	0.0005	MAVN	V	56310.7416	0.0008	SO30	BV		
	56340.9529	0.0005	HMBW	V	56310.8622	0.0008	SO30	BV		
V336 Boo	56584.2339	0.0007	AA	C	56310.9827	0.0008	SO30	BV		
	56584.2731	0.0005	AA	C	V593 Lyr	56451.4583	0.0017	AB	C	
	56584.3120	0.0004	AA	C	56478.4267	0.0007	RP	V		
	56584.3509	0.0005	AA	C	56478.5296	0.0005	RP	V		
	56584.3900	0.0006	AA	C	V1162 Ori	56355.3484	0.0022	HO40	V	
	56584.4289	0.0004	AA	C	56356.2932	0.0025	HO40	V		
	56584.4681	0.0003	AA	C	56356.3731	0.0027	HO40	V		
	56584.5073	0.0003	AA	C	56636.4933	0.0042	HO40	V		
	56584.5466	0.0005	AA	C	56637.5167	0.0027	HO40	V		
	56356.4964	0.0011	FN	C	56638.3805	0.0028	HO40	V		
V376 Cam	56403.5975	0.0004	IS	V	56638.4612	0.0032	HO40	V		
	56568.4741	0.0005	IS	V	56638.5386	0.0035	HO40	V		
	56568.6147	0.0004	IS	V	56639.4055	0.0040	HO40	V		
	56656.3146	0.0003	MAV	V	56639.4817	0.0042	HO40	V		
V435 Cam	56589.4995	0.0010	AA	C	56639.5635	0.0029	HO40	V		
AD CMi	56349.6838	0.0009	HMBW	V	56654.6679	0.0022	SO30	V		
	56349.8067	0.0011	HMBW	V	56654.7476	0.0032	SO30	V		

Table 2: Observed times of maximum (continued).

Star	Epoch	Unc.	Obs.	Filter	Star	Epoch	Unc.	Obs.	Filter		
V1162 Ori	56654.8262	0.0022	SO30	V	GSC 0632-0812	55851.7790	0.0007	HMBC	V		
	56654.9084	0.0034	SO30	V		55851.8627	0.0008	HMBC	V		
V465 Peg	56560.3500	0.0007	AB	C		55852.7904	0.0005	HMBC	V		
V536 Peg	56567.4058	0.0003	AB	C		55852.8747	0.0007	HMBC	V		
DW Psc	56558.8889	0.0003	HMBC	V		55855.7406	0.0006	HMBC	V		
	56560.8568	0.0004	HMBC	V		55855.8254	0.0010	HMBC	V		
	56564.8539	0.0002	HMBC	V	GSC 0753-1489	56341.3893	0.0015	AB	C		
56637.2662	0.0004	HO40	C	56637.4594		0.0008	AB	C			
AX Tri	56591.2745	0.0004	AA	C		56656.3897	0.0005	AB	C		
	56591.3434	0.0005	AA	C	GSC 0914-0684	56414.3073	0.0010	AA	C		
	56591.4122	0.0004	AA	C		56414.3847	0.0007	AA	C		
	56591.4812	0.0006	AA	C		56414.4623	0.0009	AA	C		
	56591.5501	0.0005	AA	C		56414.5399	0.0008	AA	C		
	56599.2081	0.0005	AA	C	GSC 0933-0651	56341.9448	0.0006	HMBW	V		
	56599.2769	0.0006	AA	C		56460.3215	0.0013	AA	C		
	56599.3458	0.0006	AA	C		56460.4259	0.0007	AA	C		
	56599.4150	0.0005	AA	C	GSC 1061-1651	56503.3033	0.0008	AA	C		
	56599.4838	0.0004	AA	C		56503.4396	0.0004	AA	C		
GW UMa	56379.5664	0.0007	IS	V	GSC 1076-0158	56458.4368	0.0006	AB	C		
	56386.4748	0.0006	IS	V		56483.4509	0.0010	AB	C		
YZ UMi	56389.4098	0.0004	VWSR	V		56528.3585	0.0047	AB	C		
	56541.3789	0.0004	VWS	V		56533.3987	0.0008	AB	C		
GSC 0191-1282	56309.4144	0.0010	IS	V	GSC 1220-1131	56320.2791	0.0007	FN	V		
	56357.3552	0.0006	FN	C		56320.3601	0.0009	FN	V		
	56357.4016	0.0006	FN	C		56337.2794	0.0016	FN	C		
	56357.4487	0.0005	FN	C		56578.3817	0.0007	AA	C		
	56366.3632	0.0003	IS	C		56578.4628	0.0007	AA	C		
	56384.3344	0.0007	FN	C		56605.2256	0.0007	AA	C		
	56384.3826	0.0006	FN	C		56605.3065	0.0006	AA	C		
	56384.4301	0.0006	FN	C		56605.3878	0.0004	AA	C		
	56384.4773	0.0006	FN	C		56605.4693	0.0004	AA	C		
	56638.4485	0.0005	AB	C		56622.4697	0.0019	RP	V		
	56638.4963	0.0003	AB	C		56646.3853	0.0009	FN	C		
	GSC 0321-0314	56364.8096	0.0008	HMBW		V	GSC 1306-0466	56309.3121	0.0036	FNA	V
	GSC 0435-3806	56461.3768	0.0013	AA		C		56366.3820	0.0014	FN	C
		56461.4407	0.0010	AA		C		56639.3443	0.0009	HO40	V
		56461.5063	0.0009	AA		C	56639.4307	0.0010	HO40	V	
GSC 0513-0624	56460.8962	0.0007	HMBC	V		56656.3212	0.0019	AB	C		
	56461.9095	0.0007	HMBC	V	GSC 1442-1358	56354.8875	0.0004	HMBW	V		
	56540.3876	0.0005	AB	C		56358.8290	0.0006	HMBW	V		
	56573.3272	0.0007	SO40	C	GSC 1594-2234	56470.4074	0.0012	SK	V		
GSC 0632-0812	55829.6929	0.0005	HMBC	V		56470.5470	0.0007	SK	V		
	55829.7762	0.0008	HMBC	V	56494.3294	0.0007	AA	C			
	55829.8617	0.0008	HMBC	V	56494.4662	0.0006	AA	C			
	55830.7895	0.0010	HMBC	V	56541.3488	0.0006	AA	C			
	55830.8728	0.0005	HMBC	V	56543.4000	0.0005	AA	C			
	55831.8002	0.0007	HMBC	V	GSC 1716-1598	56571.5260	0.0009	RP	V		
	55831.8844	0.0006	HMBC	V		GSC 1750-1237	56637.2991	0.0011	AB	C	
	55833.8230	0.0008	HMBC	V	GSC 2043-1201	56378.5892	0.0018	IS	V		
	55837.7013	0.0006	HMBC	V		56378.6677	0.0013	IS	V		
	55837.7850	0.0005	HMBC	V		56402.5160	0.0017	IS	V		
	55839.8081	0.0008	HMBC	V		56402.5933	0.0024	IS	V		
	55841.8315	0.0007	HMBC	V		56412.5688	0.0010	RP	C		
	55843.7707	0.0006	HMBC	V	GSC 2080-0986	56448.4316	0.0004	AB	C		
	55843.8551	0.0009	HMBC	V		GSC 2194-2001	56493.3230	0.0007	AA	C	
	55850.7670	0.0005	HMBC	V	56493.4375		0.0012	AA	C		
55850.8512	0.0008	HMBC	V	56552.3756	0.0011		AA	C			

Table 2: Observed times of maximum (continued).

Star	Epoch	Unc.	Obs.	Filter	Star	Epoch	Unc.	Obs.	Filter
GSC 2194-2001	56552.4907	0.0014	AA	C	GSC 2977-0238	56310.8165	0.0006	SO40	BV
	56554.3287	0.0012	AA	C		56310.8922	0.0006	SO40	BV
	56554.4431	0.0016	AA	C		56310.9682	0.0006	SO40	BV
	56555.3627	0.0013	AA	C		56390.3188	0.0011	FN	C
	56555.4774	0.0013	AA	C		56390.3948	0.0006	FN	C
	56589.2553	0.0008	AA	C		56390.4708	0.0009	FN	C
GSC 2290-1195	56589.3692	0.0012	AA	C	GSC 3004-0870	56337.3446	0.0006	AB	C
	56638.2632	0.0025	MAV	V		56365.8537	0.0008	HMBW	V
56638.3413	0.0031	MAV	V	56385.4079		0.0007	FN	C	
GSC 2496-0118	56305.4470	0.0020	IS	V	56385.4901	0.0007	FN	C	
	56337.3564	0.0029	IS	V	56385.5724	0.0007	FN	C	
	56337.4255	0.0031	IS	V	56385.6540	0.0008	FN	C	
	56379.3635	0.0006	FN	C	56400.2784	0.0005	AA	C	
	56379.4309	0.0008	FN	C	56400.3605	0.0005	AA	C	
	56379.4990	0.0010	FN	C	56400.4429	0.0005	AA	C	
	56379.5664	0.0009	FN	C	56400.5247	0.0009	AA	C	
	56379.6333	0.0013	FN	C	56416.3819	0.0007	FN	C	
	56395.2830	0.0003	AA	C	56416.4642	0.0009	FN	C	
	56395.3504	0.0004	AA	C	56416.5464	0.0011	FN	C	
	56395.4177	0.0003	AA	C	GSC 3031-0307	56356.4387	0.0006	AB	C
	56403.4143	0.0007	FN	C		56411.2798	0.0013	AA	C
	56403.4825	0.0009	FN	C		56412.2777	0.0013	AA	C
	56403.5501	0.0014	FN	C		56412.3794	0.0010	AA	C
	56638.4419	0.0009	MAV	V		56412.4785	0.0009	AA	C
	56638.5092	0.0006	MAV	V		56414.4764	0.0012	RP	C
	GSC 2566-1398	56384.6074	0.0009	IS	V	56414.5745	0.0016	RP	C
	GSC 2696-1396	56481.4575	0.0005	MAV	V	56439.3485	0.0014	AA	C
GSC 2843-1999	56254.2503	0.0007	MAV	V	56439.4491	0.0013	AA	C	
	56254.3122	0.0013	MAV	V	56447.3395	0.0011	AA	C	
GSC 2861-0970	56307.3960	0.0006	FNA	V	56447.4391	0.0010	AA	C	
	56335.3628	0.0007	FN	C	56447.5399	0.0011	AA	C	
	56573.3003	0.0013	MAV	V	GSC 3428-1497	56323.3263	0.0009	MAV	V
	56573.4100	0.0008	MAV	V		56323.4025	0.0012	MAV	V
	56577.4843	0.0006	AA	C	56379.3857	0.0018	IS	V	
	56579.3563	0.0006	AA	C	GSC 3489-0868	56349.8601	0.0005	HMBW	V
	56579.4663	0.0006	AA	C		56349.9467	0.0004	HMBW	V
	56579.5764	0.0004	AA	C		56385.3868	0.0010	MAV	V
	56592.3488	0.0007	AA	C	56385.4731	0.0005	MAV	V	
	56592.4588	0.0006	AA	C	GSC 3755-0845	56309.2795	0.0011	MAV	V
	56592.5689	0.0005	AA	C		56315.2901	0.0016	MAV	V
	56637.2714	0.0006	MAV	V		56315.3665	0.0019	MAV	V
56637.2725	0.0012	FN	C	56638.3258		0.0006	HO40	V	
56637.3814	0.0014	FN	C	56638.4020		0.0006	HO40	V	
56637.3815	0.0005	MAV	V	56638.4779		0.0005	HO40	V	
GSC 2977-0238	56299.6544	0.0004	SO40	V	56638.5541	0.0005	HO40	V	
	56299.7306	0.0005	SO40	V	56639.3150	0.0008	MAV	V	
	56299.8066	0.0006	SO40	V	GSC 3810-1553	56349.8406	0.0003	HMBW	V
	56299.8817	0.0009	SO40	V		56351.8206	0.0003	HMBW	V
	56299.9584	0.0003	SO40	V		56351.8914	0.0003	HMBW	V
	56305.5773	0.0006	IS	V	56389.3690	0.0006	FN	C	
	56309.6776	0.0008	SO40	BV	56389.4394	0.0012	FN	C	
	56309.7533	0.0006	SO40	BV	GSC 3832-0152	56308.6560	0.0012	IS	V
	56309.8293	0.0004	SO40	BV		56308.7475	0.0005	IS	V
	56309.9052	0.0006	SO40	BV		56336.6076	0.0011	IS	V
	56309.9810	0.0007	SO40	BV		56336.6984	0.0009	IS	V
	56310.6644	0.0006	SO40	BV		56378.3493	0.0003	MAV	V
56310.7405	0.0006	SO40	BV	GSC 3850-0137		56413.2733	0.0003	AA	C

Table 2: Observed times of maximum (continued).

Star	Epoch	Unc.	Obs.	Filter	Star	Epoch	Unc.	Obs.	Filter
GSC 3850-0137	56413.3366	0.0004	AA	C	GSC 4552-1498	56356.5917	0.0003	IS	V
	56413.3999	0.0004	AA	C		56356.6476	0.0003	IS	V
	56413.4633	0.0004	AA	C		56385.5020	0.0002	IS	V
	56413.5265	0.0004	AA	C		56385.5576	0.0003	IS	V
GSC 3863-0740	56378.4134	0.0014	MAVN	V		56506.4449	0.0006	MAV	V
GSC 3934-1904	56489.4786	0.0003	MAV	V	GSC 4556-1113	56306.3133	0.0004	VWSR	V
	56496.3620	0.0005	SK	V		56371.3312	0.0007	MAV	V
	56496.4712	0.0006	SK	V		56371.4173	0.0004	MAV	V
	56496.5799	0.0008	SK	V		56384.3690	0.0003	VWSR	V
GSC 4163-0984	56409.2712	0.0003	AA	C		56410.2718	0.0006	AA	C
	56409.3505	0.0005	AA	C		56410.3577	0.0008	AA	C
	56409.4300	0.0003	AA	C		56410.4446	0.0013	AA	C
	56409.5095	0.0003	AA	C		56410.5304	0.0007	AA	C
GSC 4196-1784	56463.3685	0.0012	AA	C		56487.4614	0.0004	VWSR	V
	56463.5395	0.0008	AA	C		56494.4553	0.0005	VWSR	V
GSC 4417-0394	56384.4256	0.0007	AB	C		56538.4052	0.0003	VWS	V
GSC 4464-0924	56398.5990	0.0007	IS	V		56559.2999	0.0005	VWS	V
	56485.4400	0.0009	MAV	V		56559.3861	0.0003	VWS	V
	56485.5182	0.0015	MAV	V		56637.3526	0.0004	VWS	V
	56581.3087	0.0006	IS	V	GSC 4923-0693	56353.8649	0.0004	HMBW	V
	56581.3895	0.0022	IS	V	GSC 5018-1085	56385.6112	0.0014	IS	V
GSC 4500-0083	56496.4233	0.0036	MAV	V		56418.5891	0.0014	IS	V
GSC 4552-1498	56305.6922	0.0004	IS	V	NSVS 14243430	56489.4929	0.0004	AB	C
	56336.4999	0.0004	IS	V		56539.4189	0.0004	AB	C

Table 3: Updated linear elements for some HADS. Uncertainties are given in units of the last decimal.

Star	Max (HJD)	Period (d)
GSC 1061-1651	2452383.172(7)	0.13693601(3)
GSC 1220-1131	2452625.8155(4)	0.081343587(9)
GSC 1442-1358	2452638.0670(10)	0.08211246(3)
NSVS 14243430	2452206.7164(9)	0.08607561(3)

Table 4: Independent frequencies detected in multiperiodic HADS. Uncertainties are given in units of the last decimal. The phase is given with respect to HJD = 2450000. For radial double-mode pulsators the period ratio is also listed.

Star	Frequency c/d	Semi-Amplitude Mag.	Phase	Period ratio
V879 Her	f_0 17.576979(4)	0.197(5)	0.499(4)	
	f_1 22.66134(2)	0.034(5)	0.51(3)	0.77564
GSC 1489-0914	f_0 18.453211(2)	0.120(2)	0.475(2)	
	f_1 23.860353(10)	0.0283(2)	0.175(9)	0.77338
GSC 1566-2802	f_0 16.538388(2)	0.1926(10)	0.5495(10)	
	f_1 21.97807(3)	0.0161(12)	0.2878(10)	0.75250
GSC 4145-0919	f_0 15.614244(5)	0.189(5)	0.276(4)	
	f_1 20.15552(3)	0.033(5)	0.16(3)	0.77469